

Boostrapping Gravity Solutions in 2+1 Dimensions

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Action and axi-symmetric ansatz

$$S = \frac{1}{16\pi G} \int d^3x \sqrt{-g} \left[R + 2 - \frac{1}{2}(\nabla\phi)^2 - V(\phi) \right] \quad (1)$$

Ansatz (axi-symmetric solutions):

$$ds^2 = \lambda_{\mu\nu} dx^\mu dx^\nu + \frac{e^2 d\rho^2}{-\det\lambda} \quad (2)$$

$$\lambda_{\mu\nu} = \begin{pmatrix} T+X & Y \\ Y & T-X \end{pmatrix} \quad (3)$$

$$\mathbf{X} = (T, X, Y) \quad (4)$$

$$\mathbf{X}^2 = -\det\lambda \quad (5)$$

$$S = \text{Vol} \int_{\rho_0}^{\infty} d\rho e \left[\frac{e^{-2}}{2} \dot{\mathbf{X}}^2 + 2 - \frac{e^{-2}}{2} \mathbf{X}^2 \dot{\phi}^2 - V(\phi) \right] \quad (6)$$

All variables depend on ρ only.

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Equations of motion

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 \text{Einstein:} & \quad \ddot{\mathbf{X}} = -\dot{\phi}^2 \mathbf{X} \\
 \text{Klein-Gordon:} & \quad \mathbf{X}^2 \ddot{\phi} + 2\mathbf{X} \cdot \dot{\mathbf{X}} \dot{\phi} = V'(\phi) \\
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- First-integral:

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Algorithm

Of the $SL(2, \mathbb{R})$ use two transformations to fix $\mathbf{X}(\rho)$ as

$$\mathbf{X}(\rho) = \begin{pmatrix} T \\ X \\ Y \end{pmatrix} = \frac{1}{2} \begin{pmatrix} f_1 - f_2 \\ -f_1 - f_2 \\ 0 \end{pmatrix} \quad (11)$$

In these variables, our equations become

$$\begin{array}{ll} \text{Einstein:} & \ddot{f}_{1,2} = -\phi^2 f_{1,2} \\ \text{Klein-Gordon (integrated potential):} & V(\rho) = 2 - \frac{1}{2} \frac{d}{d\rho} (f_1 \dot{f}_2) \\ \text{First-integral:} & \dot{f}_2 f_1 - \dot{f}_1 f_2 = 2j \end{array} \quad (12)$$

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- Pick $f_1(\rho)$
- Obtain f_2 by integrating the first integral: $f_2 = f_1(\gamma + 2j \int^\rho \frac{d\rho'}{f_1^2})$
- $\dot{\phi}(\rho) = \sqrt{-\ddot{f}_1/f_1}$
- $V(\phi)$ implicitly defined to satisfy K-G equation
- Plugging $f_{1,2}$ into the line element and identify the angular and time coordinates (t, φ) as linear combinations of (x^0, x^1) :

$$ds^2 = -f_2(\rho)(dx^0)^2 + f_1(\rho)(dx^1)^2 + \frac{d\rho^2}{f_1(\rho)f_2(\rho)}$$

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Static:

$$ds^2 = -f_2(\rho)dt^2 + f_1(\rho)d\varphi^2 + \frac{d\rho^2}{f_1(\rho)f_2(\rho)} \quad (13)$$

Naturally, the qualitative properties of the solutions will depend on the choice of f_1

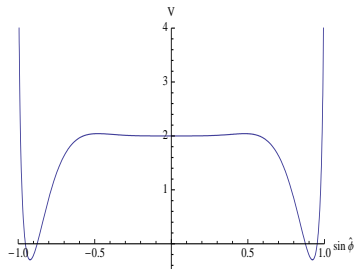
- $f_{1,2}$ linear in ρ implies asymptotically AdS space; if $f_{1,2}$ have opposite signs, the space is asymptotically dS.; f_1 linear in ρ and f_2 constant implies asymptotically flat space.
- $R = -\frac{3}{2}\dot{f}_1\dot{f}_2 - f_1\ddot{f}_2 - f_2\ddot{f}_1$. Poles may cause curvature singularities at $\rho = \rho_s$. This maybe be acceptable if the space also contains a Killing horizon or a center at $\rho_0 > \rho_s$.
- $f_2 = f_1(\gamma + 2j \int^\rho \frac{d\rho'}{f_1^2})$. If f_2 has a zero at some ρ_0 and f_1 doesn't, we have a Killing horizon.
- If f_1 has a zero at some ρ_0 and f_2 doesn't, we have a center. Furthermore we demand $f_2\dot{f}_1^2 \Big|_{\rho=\rho_0} = 4$.
- If both $f_{1,2}$ have a zero at the same point, the horizon is a Poincaré patch horizon.

Example: Asymptotically flat black holes with scalar hair

Remember: $ds^2 = -f_1(\rho)dt^2 + f_2(\rho)dx^2 + \frac{d\rho^2}{f_1(\rho)f_2(\rho)}$

Solution:

- $f_1 = 1 - \frac{1}{\rho}$, $\rho \in (1, \infty)$, $\dot{f}_2 f_1 - \dot{f}_1 f_2 = 2j$
- $f_2 = f_1 \left(A - 2j \left[\frac{1}{\rho-1} - \rho - 2 \log(\rho-1) \right] \right) \simeq 2j\rho + \dots$
- $R = \frac{5j}{\rho^2} + \dots$
- $\phi = -\sqrt{2}\pi + 2\sqrt{2} \arctan \sqrt{\frac{\rho}{\Omega} - 1} \simeq -2\sqrt{2} \sqrt{\frac{\Omega}{\rho}} + \dots$, $\phi \in (-\pi\sqrt{2}, 0)$
- $V(\phi) = 2 + jx^4 + x^6(A - 4j + 4j \log \frac{1-x^2}{x^2}) - \frac{3}{2}x^8(A + 2j + 4j \log \frac{1-x^2}{x^2})$, $x = \sin \frac{\phi}{2\sqrt{2}}$



Can we find a family of solutions?

$$\begin{array}{llll}
 \text{Einstein:} & \ddot{\mathbf{X}} = -\dot{\phi}^2 \mathbf{X} & \rightarrow +6 & \\
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We used two transformations of $SL(2, \mathbb{R})$ to write

$$\mathbf{X}(\rho) = \frac{1}{2} \begin{pmatrix} f_1 - f_2 \\ -f_1 - f_2 \\ 0 \end{pmatrix} \rightarrow -2 \tag{15}$$

Of the remaining **five** integration constants, **three** are gravity

- Lorentz boosts $f_1 \rightarrow \gamma f_1$, $f_2 \rightarrow f_2/\gamma$
- Shifts $\rho \rightarrow \rho + \rho_0$
- Particle angular momentum j

and the other **two** are the scalar leading and subleading modes ϕ_{\pm} .

We have a total of **five** parameters: $\gamma, \rho_0, j, \phi_{\pm}$.

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Family of solutions

Rescalings!

$$\rho \rightarrow \Omega\rho \quad f_{1,2} \rightarrow \Omega f_{1,2} \quad j \rightarrow \Omega j \quad (16)$$

$$\frac{d}{d\rho} \rightarrow \frac{1}{\Omega} \frac{d}{d\rho} \quad \phi \rightarrow \phi \quad (17)$$

$$\dot{f}_{1,2} \rightarrow \dot{f}_{1,2} \quad \ddot{f}_{1,2} \rightarrow \frac{1}{\Omega} \ddot{f}_{1,2} \quad (18)$$

$$\begin{aligned} \text{Einstein:} & \quad \ddot{f}_{1,2} = -\dot{\phi}^2 f_{1,2} \\ \text{Klein-Gordon (integrated potential):} & \quad V(\rho) = 2 - \frac{1}{2} \frac{d}{d\rho} (f_1 \dot{f}_2) \\ \text{First-integral:} & \quad \dot{f}_2 f_1 - \dot{f}_1 f_2 = 2j \end{aligned} \quad (19)$$

Example: Asymptotically flat black holes - family of solutions

Remember: $ds^2 = -f_2(\rho)dt^2 + f_1(\rho)dx^2 + \frac{d\rho^2}{f_1(\rho)f_2(\rho)}$

Solution:

- $f_1 = 1 - \frac{\Omega}{\rho}$, $\Omega > 0$, $\rho \in (\rho_h, \infty)$, $f_2(\rho_h) = 0$, $\rho_h > \Omega$
- $f_2 = \Omega^2 f_1 \left(A - 2\xi \left[\frac{\Omega}{\rho - \Omega} - \frac{\rho}{\Omega} - 2 \log\left(\frac{\rho}{\Omega} - 1\right) \right] \right)$ $\dot{f}_2 f_1 - \dot{f}_1 f_2 = 2\xi\Omega$
- $R = 5\xi \frac{\Omega^2}{\rho^2} + \dots$
- $\phi = -\sqrt{2}\pi + 2\sqrt{2} \arctan \sqrt{\frac{\rho}{\Omega} - 1} \simeq -2\sqrt{2} \sqrt{\frac{\Omega}{\rho}} + \dots$, $\phi \in (-\pi\sqrt{2}, 0)$
- $V(\phi) = 2 + \xi x^4 + \left(A - 4\xi + 4\xi \log \frac{1-x^2}{x^2} \right) x^6 - \frac{3}{2} \left(A + 2\xi + 4\xi \log \frac{1-x^2}{x^2} \right) x^8$
 $x = \sin(\phi/\sqrt{8})$

Asymptotics

Brown-Henneaux asymptotics:

$$f_{1,2} = 2\rho + B_{1,2} + \mathcal{O}(\rho^{-\delta_{1,2}}), \quad 1 \leq \delta_{1,2} > 0 \quad (20)$$

Simple generalization of Brown-Henneaux asymptotics:

$$f_1 = 2\rho + A\Omega^{1-\epsilon}\rho^\epsilon + B\Omega \log \rho + \rho_0 + \dots \quad (21)$$

$$\phi = \phi_- \rho^{-\Delta_- / 2} + \phi_+ \rho^{-\Delta_+ / 2} + \dots, \quad \Delta_\pm = 1 \pm \epsilon, \quad m^2 = \epsilon^2 - 1 \quad (22)$$

Holographic renormalization and thermodynamics

$$\Gamma_+ = \int d^3x \sqrt{-g} \left[R + 2 - \frac{1}{2}(\partial\phi)^2 - V(\phi) \right] + 2 \int d^2x \sqrt{-\gamma} [K - U(\phi)]$$

$$S_{\mathcal{W}} = -4\text{Vol}\epsilon\phi_-\phi_+ - 2\text{Vol}(\mathcal{W}(\phi_-) - \phi_-\mathcal{W}'(\phi_-)) \quad (23)$$

$$U(\phi) = 1 + \frac{1-\epsilon}{4}\phi^2 + \dots \quad (24)$$

Dirichlet	$\mathcal{J} = \phi_-$	(25)
Neumann	$\mathcal{J} = 2\epsilon\phi_+$	
Mixed	$\mathcal{J} = 2\epsilon\phi_+ - \mathcal{W}'(\phi_-)$	

$$\langle T_{ij} \rangle_+ = -2 \lim_{\rho \rightarrow \infty} [K_{ij} - (K - U(\phi))\gamma_{ij}] \quad (26)$$

$$\langle T_{ij} \rangle_- = \langle T_{ij} \rangle_+ - (\mathcal{W}(\phi_-) + \phi_-\mathcal{J})g_{(0)\mu\nu} \quad (27)$$

Also, thermodynamics: temperature, mass, angular momentum and entropy

Solutions

- Reproduced results from the literature: axi-symmetric solutions
- Brown-Henneaux solitons
- Black holes (BH and non-BH)
- Solutions with exponential potential (neither flat nor AdS)